

Extended Abstract

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Orlando, Florida, June 2003**Graphite Ablation and Thermal Response Simulation
Under Arc-jet Flow Conditions**Y.-K. Chen, F. S. Milos, D. C. Reda, and D. A. Stewart
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Moffett Field, CA 94035-1000**Abstract**

The Two-dimensional Implicit Thermal Response and Ablation program, TITAN, was developed and integrated with a Navier-Stokes solver, GIANTS, for multi-dimensional ablation and shape change simulation of thermal protection systems in hypersonic flow environments. The governing equations in both codes are discretized using the same finite-volume approximation with a general body-fitted coordinate system. Time-dependent solutions are achieved by an implicit time marching technique using Gauss-Siedel line relaxation with alternating sweeps.

As the first part of a code validation study, this paper compares TITAN-GIANTS predictions with thermal response and recession data obtained from arc-jet tests recently conducted in the Interaction Heating Facility (IHF) at NASA Ames Research Center. The test models are graphite sphere-cones. Graphite was selected as a test material to minimize the uncertainties from material properties. Recession and thermal response data were obtained from two separate arc-jet test series. The first series was at a heat flux where graphite ablation is mainly due to sublimation, and the second series was at a relatively low heat flux where recession is the result of diffusion-controlled oxidation. Ablation and thermal response solutions for both sets of conditions, as calculated by TITAN-GIANTS, are presented and discussed in detail. Predicted shape change and temperature histories generally agree well with the data obtained from the arc-jet tests.

NomenclatureA = area, m² $B' = \dot{m} / \rho_e u_e C_m$, dimensionless mass blowing rate B_a = pre-exponential constant in Eq.(6), s⁻¹ C_h = Stanton number for heat transfer C_m = Stanton number for mass transfer c_p = specific heat, J/kg-KE = total energy per unit volume, J/m³ E_a = activation energy in Eq.(6), J/kmol

F = radiation view factor in Eq.(11)

h = enthalpy, J/kg

 \bar{h} = enthalpy of the partial-pyrolyzed solid, defined in Eq.(4), J/kg

h_j^{Tw} = enthalpy of species j respect to surface temperature in Eq.(11), J/kg
 H_r = recovery enthalpy, J/kg
 k = thermal conductivity, W/m-K
 \dot{m} = mass flux, kg/m²-s
 p = pressure, N/m²
 q_C = conductive heat flux, W/m²
 q_R = radiative heat flux, W/m²
 q_v = heat flux due to species diffusion, W/m²
 R = universal gas constant, J/kmol-K
 R_n = nose radius, m
 T = temperature, K
 t = time, s
 u = fluid velocity, m/s
 v = local grid velocity, m/s
 x = Cartesian coordinate system, m
 Z^* = coefficient in Eq.(11), defined in Ref. 14
 α = surface absorptance
 ϵ = surface emissivity
 Γ = volume fraction of resin
 λ = blowing reduction parameter in Eq.(12)
 ρ = total density, kg/m³
 ρ_o = original density, kg/m³
 ρ_r = residual density, kg/m³
 σ = Stefan-Boltzmann constant, W/m²-K⁴
 τ = mass fraction of virgin material defined in Eq.(3), or shear stress in Eq.(9)
 Ψ = decomposition reaction order in Eq.(6)
 ξ, η = general body fitted coordinate system defined in Fig.2
 v = species diffusion flux, m/s
 w = species source term in Eq.(8), kg/m³-s
 ∇ = gradient, m⁻¹
subscripts
 c = char
 e = boundary-layer edge
 g = pyrolysis gas
 i = material density component (A, B, and C), or direction component
 j = surface species, or direction component
 s = gas species or stagnation point
 v = virgin
 w = wall

Introduction

Heatshields of spacecraft for planetary mission typically use thermal protection system (TPS) materials which pyrolyze and ablate at high temperature for mass-efficient rejection of aerothermal

heat load. Pyrolysis is an internal decomposition of the solid which releases gaseous species, whereas ablation is a combination of processes which consume heatshield surface material. A reliable numerical procedure that can compute surface recession rate, mass loss, in-depth pyrolysis, and internal temperature time histories under general heating conditions is essential for the design and sizing of ablating spacecraft TPS materials.

The purpose of this paper is to compare TITAN-GIANTS predictions with arc-jet data and to study the accuracy of the simulation. Thermal response and recession data were obtained from arc-jet tests recently conducted in the Interaction Heating Facility (IHF) at NASA Ames Research Center. Graphite ablation and thermal properties have been widely studied, so graphite was selected as the test material to minimize the uncertainties from material properties (the accuracy of shape change and temperature history predictions can thus be better determined). However, the modeling of pyrolysis gas cannot be examined from graphite data. The pyrolysis gas effect will be studied in future work.

The recession and thermal response data studied here are from two separate arc-jet test series. The first series was at relatively high heat flux in which graphite ablation was mainly due to sublimation, and the second series was at relatively low heat flux in which the recession was the result of diffusion-controlled oxidation. Typical B' vs. temperature curves for graphite in air at various surface pressures (0.01, 0.1, and 1 atm) are shown in Fig. 1. Above 3000 K, B' is a strong function of temperature, and sublimation is the primary ablation process. At lower temperature, B' is relatively constant (approximately 0.175), and diffusion-controlled oxidation determines the ablation rate. At temperatures below 1100 K, the ablation is kinetically controlled and is beyond the scope of this work.

Background

The Charring Material Thermal Response and Ablation Program¹ (CMA) was developed by Aerotherm Corporation in the 1960s. This technique solved the one-dimensional internal energy balance and decomposition equations coupled with the ablating surface energy balance condition to simulate the response of ablative heatshields in hypersonic flows. The Fully Implicit Ablation and Thermal Response program² (FIAT) was developed at NASA Ames Research Center in the 1990s. Both FIAT and CMA solve the same governing equations. FIAT, which is numerically more stable, solves a wider range of problems compared with CMA, and has been used for TPS sizing calculations for various NASA space missions, such as Mars Pathfinder³ and Stardust.⁴

The Two-dimensional Implicit Thermal Response and Ablation (TITAN) program⁵ was developed and coupled with the GIANTS (Gauss-Seidel implicit aerothermodynamic Navier-Stokes code with thermochemical surface boundary conditions) code⁶⁻⁹ using a loosely coupled method to perform high fidelity thermal response and shape change simulation for charring materials. The governing equations in both TITAN and GIANTS are discretized using the same finite-volume approximation with general body fitted coordinates. The solution is achieved by an implicit time marching technique using Gauss-Seidel line relaxation with alternating sweeps. A fluid-solid coupled simulation is required to accurately predict shape change and aerothermal

heating, because aerothermal heating distribution is very sensitive to body shape, and shape change is a strong function of surface heating distribution. TITAN-GIANTS was also integrated with the commercial finite-element code, MARC,¹⁰ to perform thermal and structural analysis on the complex geometries of actual vehicles.¹¹ Details of TITAN's development and its formulation have been discussed in a previous paper.⁵ Application of TITAN-GIANTS and TITAN-MARC for full body heat-shield/structure analysis of a Mars sample return Earth Entry Vehicle and for simulation of a flat-faced cylinder arc-jet test model are demonstrated in Reference 11.

Arc-Jet Facilities

Arc-jet facilities are used to simulate the aerothermal heating environments experienced by space vehicles during atmospheric entry. The arc-jet tests discussed in this paper were conducted in the Ames 60 MW Interaction Heating Facility. In the IHF, air is heated by the electric discharge in an 8-cm diameter, water-cooled constricted arc column. The hot gas is then expanded through a converging-diverging nozzle into an evacuated test chamber where the test model is located. The heat flux and pressure on the model surface can be varied by changing the electrical power input and the configuration of the nozzle, as well as the model location from the nozzle exit. The test model is mounted onto a mechanical swing arm and positioned between four and fourteen inches from the nozzle exit. The swing arm inserts the test model into the gas stream when the desired test conditions are reached, and then retracts the model from the stream after the desired exposure time has elapsed.

Two-dimensional Implicit Thermal and Ablation Program – TITAN

The ablating heat-shield material thermal response and shape change computations are performed using TITAN. The governing equations include energy conservation and a three-component decomposition model. The surface energy balance condition is solved with a moving grid to calculate the shape change due to surface recession. The internal energy balance is a transient thermal conduction equation with additional pyrolysis terms⁵

$$\rho c_p \frac{\partial T}{\partial t} = \nabla \cdot (k \nabla T) - (h_g - \bar{h}) \nabla \cdot \dot{m}_g + \dot{m}_g \cdot \nabla h_g + v \rho c_p \nabla T \quad (1)$$

The individual terms in Eq. 1 may be interpreted as follows: rate of storage of sensible energy, net rate of thermal conductive heat flux, pyrolysis energy-consumption rate, net rate of energy convected by pyrolysis, and convection rate of sensible energy due to coordinate system movement. The local specific heat is formulated from functions of temperature input for both virgin material and char. In partially pyrolyzed zones ($\rho_c < \rho < \rho_v$), the specific heat is obtained from the mixing rule

$$c_p = \tau c_{pv} + (1 - \tau) c_{pc}, \quad (2)$$

where τ is the mass fraction of virgin material in a mixture of virgin material and char:

$$\tau = \frac{1 - \rho_c / \rho}{1 - \rho_c / \rho_v} . \quad (3)$$

The thermal conductivity, k , is weighted in the same manner. The pyrolysis gas enthalpy, h_g , is an input function of temperature and pressure. The enthalpy of the partial-pyrolyzed solid, \bar{h} , is defined as

$$\bar{h} = \frac{\rho_v h_v - \rho_c h_c}{\rho_v - \rho_c} . \quad (4)$$

A three-component decomposition model is used. The resin filler is assumed to consist of two components which decompose separately, while the reinforcing material is the third component which can decompose. The instantaneous density of the composite is given by:

$$\rho = \Gamma(\rho_A + \rho_B) + (1 - \Gamma)\rho_C , \quad (5)$$

where A and B represent components of the resin, and C represents the reinforcing material. Γ is the volume fraction of resin and is an input quantity. For graphite, the volume fraction, Γ , is 0, and the reinforcing material density, ρ_C , is the graphite density. Each of the three components can decompose by the relation:

$$\frac{\partial \rho_i}{\partial t} = -B_{ai} \exp\left(\frac{-E_{ai}}{RT}\right) \rho_{oi} \left(\frac{\rho_i - \rho_{ri}}{\rho_{oi}}\right)^{\Psi_i} + v \nabla \rho_i , \quad (6)$$

where ρ_{ri} is the residual or terminal density of component i , and ρ_{oi} is the original density of component i ($i = A, B$, and C). The motion of pyrolysis gas is assumed to be quasi-steady and one-dimensional (along the η direction), and thus the mass flow rate of pyrolysis gas at the surface is calculated using the following approximation:

$$\dot{m}_g = \frac{1}{A_w} \int_0^{\eta_w} -\left(\frac{\partial \rho}{\partial t}\right) A_\eta d\eta . \quad (7)$$

The governing equations are discretized using a finite-volume approximation with a general body fitted coordinate system. The time-dependent solution is achieved by a implicit time marching technique using Gauss-Siedel line relaxation with alternating sweeps. If time accuracy is not of interest in the computation, the time step should be as large as possible, and the number of alternating sweeps in each time step should be 1. A time-accurate solution can be achieved by increasing the number of alternating sweeps in each time step or by reducing the time step. The computational grid is compressed during the course of computation to reflect the surface recession. Further details of the numerical procedures are presented in Reference 5.

Navier-Stokes equation solver – GIANTS

To estimate the aerothermal heating distribution over a large angle blunt body, the hypersonic flow simulation is performed using the Navier-Stokes solver, GIANTS. The GIANTS code solves the time-dependent conservation equations of mass, momentum, and energy for the chemical and thermal non-equilibrium flow-field. The species mass conservation equation is given by¹²

$$\frac{\partial \rho_s}{\partial t} + \frac{\partial}{\partial x_j} (\rho_s u_j) = - \frac{\partial}{\partial x_j} (\rho_s v_{sj}) + w_s, \quad (8)$$

the total momentum conservation is written as

$$\frac{\partial}{\partial t} (\rho u_i) + \frac{\partial}{\partial x_j} (\rho u_i u_j) = - \frac{\partial \tau_{ij}}{\partial x_j}, \quad (9)$$

and the total energy conservation as

$$\frac{\partial E}{\partial t} + \frac{\partial}{\partial x_j} ((E + p) u_j) = - \frac{\partial}{\partial x_j} (q_j + q_{vj}) - \frac{\partial}{\partial x_j} (u_i \tau_{ij}) - \sum_{s=1}^n \frac{\partial}{\partial x_j} v_{sj} h_s. \quad (10)$$

The governing equations are discretized using the finite-volume method. The numerical method (fully implicit and Gauss-Seidel line relaxation) used to solve the discretized equations is exactly the same as that used in the TITAN code. This technique has been shown to yield steady-state results efficiently. A five species (N₂, O₂, NO, O, and N) gas model is used to accurately model the high temperature behavior of air. A bifurcation diffusion model¹³ was also implemented in this version of the code to correctly compute multi-species mass diffusion.

TITAN and GIANTS are interfaced through the surface energy balance boundary condition. The conditions at the ablating surface are determined by convective and radiative heating and by surface thermochemical interactions with the boundary layer gases. The surface energy balance equation employed is of the convective transfer coefficient type. This energy balance equation takes the following form¹

$$\begin{aligned} \rho_e u_e C_h (H_r - h_{ew}) + \rho_e u_e C_m \left[\sum (Z_{je}^* - Z_{jw}^*) h_j^{Tw} - B' h_w \right] + \dot{m}_c h_c + \dot{m}_g h_g \\ + \alpha_w q_{Rw} - F \sigma \epsilon T_w^4 - q_{Cw} = 0 \end{aligned} \quad (11)$$

The first term in Eq. 11 represents the sensible convective heat flux. The sum of the second, third, and fourth terms in Eq. 11 is defined as the total chemical energy at the surface. The Z^* terms represent transport of chemical energy associated with chemical reactions at the wall and in the boundary layer.¹⁴ The Z^* driving forces for diffusive mass transfer include the effects of unequal diffusion coefficients. The fifth and sixth terms are the radiative heat fluxes absorbed and re-radiated by the wall, respectively, and the last term, q_{Cw} , represents the rate of conduction into the material. The radiative heat flux, q_{Rw} , such as radiation from a hot shock layer, is an input quantity. The prediction of shock layer radiation is not currently included in the coupled TITAN-GIANTS simulation. Here B' is the normalized mass blowing rate. The tables of B' for charring materials can be pre-generated using ACE¹⁴ or MAT¹⁵ assuming chemical equilibrium.

The coupled TITAN-GIANTS simulation does not include surface mass blowing and ablation chemical species in the Navier-Stokes computation. Instead, a blowing

correction parameter is introduced to account for the reduction in transfer coefficients due to the transpiration effect of the mass injection into the boundary layer. This approach was taken because previous studies of heat-shield design for Mars Pathfinder³ and Stardust⁴, demonstrated that the computational time increased by about at least one order of magnitude if a blowing surface condition and ablation chemical species with their associated reactions were included in a flow simulation. In addition, the reaction rates of many gas phase chemical reactions are unknown or not well studied. Thus, the prediction of Navier-Stokes solver becomes unreliable because of the uncertainties on the gas phase chemical reaction rates. The introduction of a blowing reduction parameter can give a reasonable prediction for many applications and significantly cut the CPU time. The blowing rate correction for convective heat transfer is

$$\frac{C_h}{C_{h1}} = \frac{\ln(1 + 2\lambda B')}{2\lambda B'} \quad , \quad (12)$$

where λ is the blowing reduction parameter, C_h is the heat transfer coefficient for the ablating surface, and C_{h1} is the heat transfer coefficient for a nonablating surface. When $\lambda = 1/2$, Eq.(11) reduces to the classical laminar-flow blowing correction.¹⁶

The computation starts with the TITAN code using initial estimates of surface heat flux and pressure without shape change as the surface boundary condition. At each time increment, TITAN obtains the in-depth thermal response solution based on current boundary conditions, and also calls the flow environment code to update the front-face energy balance boundary condition if necessary. The non-ablating surface heating and pressure are calculated by GIANTS, and a blowing reduction parameter of 0.5 is used in TITAN to take into account the laminar flow blockage due to surface mass blowing. When the maximum local surface recession exceeds a predefined criterion since the last surface convective heating was updated, a flow-field grid is generated based on the current body geometry, and then the flow simulation routine is called upon to compute the new aerothermal heating environment. Each call to the flow environment routine is a steady state calculation. As expected, the GIANTS calculation is much more computationally intensive than the thermal response computation performed by TITAN. For a TITAN-GIANTS coupled simulation, most of the CPU time is consumed by the flow environment calculation. Thus, the CPU time required for a simulation is primarily determined by the efficiency of the GIANTS code.

Results and Discussion

The first series of arc-jet tests were conducted with a stream enthalpy of 27 MJ/Kg, a stagnation point cold-wall heat flux of 2100 W/cm², and a stagnation point pressure of 0.75 atm. The length of the heat pulse is 30 sec. The Poco graphite model is a 10° half angle sphere-cone with nose radius of 1.905 cm (0.75"). The thermal properties of Poco graphite are from Ref. 17. The total length of the model is 8.89 cm (3.5"). The computational grids for both solid and fluid are shown in Fig. 2, where the solid grid is divided into two zones, and the fluid grid has a single zone. The first zone in the solid will be reconstructed as the recession exceeds preset criteria, and the second zone remains unchanged through the entire computation. Surface heat flux and pressure

distributions are recalculated using GIANTS when free stream conditions or the first zone in the solid grid is changed. The thermal response computation of the solid is transient, and the flow field is assumed to reach steady state instantaneously.

The heat flux distributions over the solid surface at $t = 0$ and 30 sec are shown in Fig. 3. The stagnation point heating rate at the end of heat pulse ($t = 30$ sec) is about 90% of that at the beginning ($t = 0$ sec). However, the heating at 45° from stagnation point becomes slightly higher toward the end of the heat pulse. Surface recession tends to blunt the spherical nose, and thus reduce stagnation point heating. Because of the uncertainties in estimated arc-jet free stream conditions and material surface conditions, the GIANTS solution at $t = 0$ does not exactly match the measured stagnation-point cold-wall heat flux (2100 W/cm^2). To account for these uncertainties, the normalized heat flux distributions from GIANTS were scaled to exactly match the measured initial stagnation point heat flux.

A comparison of predicted and measured surface recession is shown in Fig. 4. The symbols for the graphite surface recession data are measured every 5 seconds at the stagnation point and at 45° from the stagnation point. (Details of the measurements are described in Ref. 18.) The dark solid lines are the predicted recession using a coupled TITAN-GIANTS simulation. The total recession for this 30-sec heat pulse at stagnation point is about 0.33 cm and at 45° from stagnation point is about 0.165 cm. The computation slightly over-predicts the recession at the stagnation point and slightly under-predicts the recession at 45° . The maximum difference between prediction and measurement is less than 10 %.

For comparison purposes, an additional FIAT computation was performed by assuming that the cold-wall heat flux was 2100 W/cm^2 and the thermal diffusion was one-dimensional. FIAT predicts much higher recession compared with TITAN at the stagnation point, as shown by the lighter solid line in Fig. 4. This result is not surprising, because the FIAT calculation does not take into account multidimensional heat conduction and shape change effects.

The B' histories at the stagnation point and at 45° are presented in Fig. 5. In the first 2 seconds, the stagnation point is in the diffusion-controlled regime, but as the surface temperature increases the ablation enters the sublimation regime, and B' levels off near 0.4 until the end of heat pulse. At 45° from the stagnation point, the surface remains in the diffusion controlled regime until about 5 seconds, and then B' gradually increases to about 0.3 at 30 sec.

The predicted surface temperature history at 45° from the stagnation point is shown in Fig. 6 and compared with temperature measurements from an infrared (IR) camera and a pyrometer. At $t = 30$ sec, the predicted surface temperature is 3180°C , the temperature measured from IR camera is 3030°C , and that from the pyrometer is 2940°C . The difference between the prediction and the IR is about 5% and between the prediction and the pyrometer is about 8 %. Further study is required to fully understand why the calculated temperature is higher than the measured data, and why the temperature reading from the IR camera is higher than that from the pyrometer.

There were no stagnation point surface temperature and in-depth temperature measurements in this test series. The predicted temperature contours of the graphite model at $t = 30$ sec are presented in Fig. 7. The stagnation point temperature is above

3500 K, and the back face of test model, which is connected to a water cooled holder, is assumed to remain at 312 K. The dark solid line represents the initial surface contour. The area between the dark solid line and color contours is the material ablated during the 30 sec heat pulse.

Figure 8 presents the flow field predictions at 0 and 30 sec. The upper half is the flow temperature contours at the beginning of test ($t = 0$ sec), and the lower half is that at the end of the heat pulse ($t = 30$ sec). The Knudsen number is about 0.023. As expected, there is no significant change in flow pattern. The maximum fluid temperature around the shock is above 10,000 K. The shock stand-off distance at $t = 30$ sec is slightly larger than at $t = 0$ sec. This is because ablation blunts the spherical section, and consequently reduces the heating in the stagnation region.

The second arc-jet test series was conducted with a stream enthalpy of 19 MJ/Kg, a stagnation point cold wall heat flux of 593 W/cm^2 , and a stagnation point pressure of 0.05 atm. The duration of heat pulse was 60 sec. The ATJ graphite composite model is a 15° half angle sphere-cone with nose radius of 0.95 cm (0.375"). The total length of model is 10.16 cm (4"). The material map and computational grids for both solid and fluid are shown in Fig. 9. The solid grid is again divided into two zones, and the fluid grid has a single zone. The first 3.81 cm (1.5") of the model is a mushroom-shaped piece of ATJ graphite.¹⁹ The rest of the model is built from a AETB 35/12 insulator with a small region of TUFROC to prevent overheating of the bond-line to the AETB.

The heat flux distributions over the solid surface at $t = 0$ and 60 sec are shown in Fig. 10. The stagnation point heating rate at the end of heat pulse ($t = 60$ sec) is about 95% of that at the beginning ($t = 0$ sec), and again (similar to the first series) at 45° from stagnation point the heat flux becomes slightly higher at the end of heat pulse. However, the difference in surface heating between the beginning and the end of heat pulse is relatively small, compared with the first test series, because the total recession is lower in this second series of tests.

The stagnation point recession, measured at $t = 25$ and 60 sec, is compared with predictions in Fig. 11. The solid line is the prediction from TITAN, and the dashed line is from FIAT (a one-dimensional calculation with constant heat flux of 593 W/cm^2). As expected, FIAT predicts slightly higher recession compared with TITAN. The maximum difference between TITAN and FIAT is about 9% at $t = 60$ sec. The measured stagnation point recessions at 25 and 60 sec are 0.068 and 0.183 cm, respectively. The predicted recession from TITAN at 25 sec and 60 sec is 0.071 cm and 0.174 cm, respectively. The maximum difference between the TITAN prediction and the measurement is about 5%. In this test series, graphite ablation is mainly due to diffusion-controlled oxidation. Thus, B' remains relatively constant (approximately 0.175) through the entire heat pulse. The total mass loss in this second test series is smaller than in the first one.

Three thermocouples (designated TC1, TC2, and TC3) are located on the model axis at depths of approximately 3.18, 4.13, and 6.67 cm from the initial stagnation point. The predicted temperature histories at these three locations from TITAN (dark solid lines) are presented and compared with the TC data in Fig. 12. Generally speaking, the TITAN predictions are in good agreement with the data. The TITAN predicted maximum temperatures of TC1 and TC2 are slightly below the data, but the predicted cooling rates for these two TC's are slightly lower than data. At the location of TC3, the TITAN prediction is slightly lower than the data most of the time.

Surface temperature data are not available for this test series. Predicted temperature contours at $t = 60$ sec (end of the heat pulse) and at 200 sec (during cooldown) are shown in Figs. 13 and 14, respectively. The peak temperature at 60 sec is above 2700 K around stagnation point, and at 200 sec the hottest location is about 1200 K near the interface between the graphite and the AETB. The back face of model is connected to a water cooled holder and assumed to be at 312 K. The dark solid line again represents the initial surface contour.

Figure 15 presents the flowfield predictions at 0 and 60 sec. The upper half is the temperature contours at the beginning of heat pulse ($t = 0$ sec), and the lower half is that at the end of heat pulse ($t = 60$ sec). The Knudsen number for this case is about 0.1. These results are similar to the first series (Fig.8) in that there is no significant change in flow pattern. The maximum fluid temperature around the shock is above 7000 K. The shock stand-off distance at $t = 60$ sec is almost the same as that at $t = 0$ sec.

Conclusions

Predictions from the integrated TITAN-GIANTS code were compared with thermal response and recession data obtained from arc-jet tests conducted in the Interaction Heating Facility at NASA Ames Research Center. The test models are graphite sphere-cones. The arc-jet data were taken from two separate test series. The first test series ($q_s = 2,100 \text{ W/cm}^2$, and $P_w = 0.75 \text{ atm.}$) had graphite ablation in the sublimation regime, and the second test series ($q_s = 593 \text{ W/cm}^2$, and $P_w = 0.05 \text{ atm.}$) had ablation in the diffusion controlled oxidation regime. Ablation and thermal response solutions calculated by TITAN-GIANTS for both sets of test conditions agree reasonably well with the measurements. The calculations slightly over-predict the surface temperature at 45° from the stagnation point for the sublimation case, and slightly under-predict in-depth maximum temperatures for diffusion-controlled case. The maximum difference in recession between TITAN predictions and measurement is less than 10 %. A one-dimensional FIAT computation can be used to roughly estimate the stagnation point recession, but not to predict in-depth thermal response.

References

- ¹Anon., *User's Manual: Aerotherm Charring Material Thermal Response and Ablation Program*, Acurex Corporation, Aerotherm Division, Mountain View, California, Aug. 1987.
- ²Chen, Y.-K., and Milos, F.S., "Ablation and Thermal Analysis Program for Spacecraft Heatshield Analysis," *J. of Spacecraft and Rockets*, Vol. 36, No. 3, May-June 1999, pp. 475-483.
- ³Chen, Y.-K., Henline, W.D., and Tauber, M.E., "Mars Pathfinder Trajectory Based Heating and Ablation Calculations", *J. of Spacecraft and Rockets*, Vol. 32, No. 2, March-April. 1995, pp 225-230.
- ⁴Olynick, D.R., Chen, Y-K, and Tauber, M.E., "Aerothermodynamics of the Stardust Sample Return Capsule", *J. of Spacecraft and Rockets*, Vol. 36, No. 3, May-June 1999, pp. 442-462.

- ⁵Chen, Y.-K., and Milos, F.S., "Two-Dimensional Implicit Ablation Thermal Response and Ablation Program for Charring Materials on Hypersonic Space Vehicles," AIAA Paper 2000-0206, January 2000.
- ⁶MacCormack, R.W., "Current Status of Numerical Solutions of the Navier-Stokes Equation," AIAA Paper 85-0032, Jan 1985.
- ⁷Candler, G.V., and MacCormack, R. W., "Hypersonic Flow Past 3-D Configuration," AIAA Paper 87-0480, June 1987.
- ⁸Candler, G.V., "The Computation of Hypersonic Ionized Flows in Chemical and Thermal Nonequilibrium," AIAA Paper 88-0511, January 1988.
- ⁹Chen, Y.-K., Henline, W.D., Stewart, D.A., and Candler, G.V., "Navier-Stokes Solutions with Surface Catalysis for Martian Atmospheric Entry," *J. of Spacecraft and Rockets*, Vol 30, No 1, Jan-Feb 1993, pp. 32-42.
- ¹⁰Anon., *Users Manual, MARC Analysis Research Corporation, Volume A: Users Information*, 1994, MARC Analysis Research Corporation, Palo Alto, CA.
- ¹¹Chen, Y.-K., and Milos, F.S., "Thermal Response Modeling System for a Mars Sample Return Vehicle", The 12th Thermal and Fluids Analysis Workshop, Sept. 2001.
- ¹²Lee, J.H., "Basic Governing Equations for the Flight Regimes of Aeroassisted Orbital Transfer Vehicles", Thermal Design of Aeroassisted Orbital Transfer Vehicles, edited by H.F.Nelson, Vol. 96, Progress in Astronautics and Aeronautics, AIAA, New York, 1985, pp. 3-53.
- ¹³Moyer, C.B., and Rindal, R.A., "Finite Difference Solution for the In-Depth Response of Charring Materials Considering Surface Chemical and Energy Balances," Aerotherm Corporation, Mountain View, California, Final Report 66-7, Part II, March 1967 (also NASA CR-1061, June 1968).
- ¹⁴Anon., *User's Manual: Aerotherm Chemical Equilibrium Computer Program*, Acurex Corporation, Aerotherm Division, Mountain View, California, Aug. 1981.
- ¹⁵Milos, F.S., and Chen, Y.-K., "Comprehensive Model for Multi-Component Ablation Thermochemistry," AIAA Paper 97-0141, Jan. 1997.
- ¹⁶Kays, W.M., and Crawford, M.E., *Convective Heat and Mass Transfer*, 2nd Edition, McGraw-Hill, 1980, pp. 355-357.
- ¹⁷Private Communication (from Al Covington)
- ¹⁸Reda, D., To be published as NASA TM.
- ¹⁹Touloukian, Y. S., Powell, R. W., Ho, C. Y., and Klemens, P. G., *Thermal Conductivity Nonmetallic Solid, Thermophysical Properties of Matter*, Vol. 2, 1970, pp. 21-23.

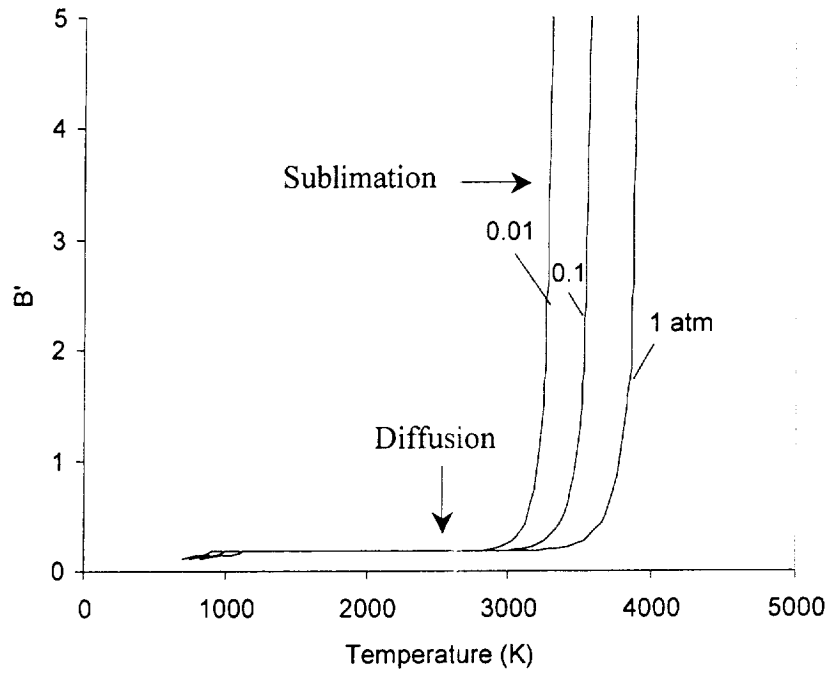


Figure 1: Graphite B' curves in air (generated using JANAF data)

$q_s = 2,100 \text{ W/cm}^2$
 $P = 0.75 \text{ atm}$
 $H = 27 \text{ MJ/Kg}$

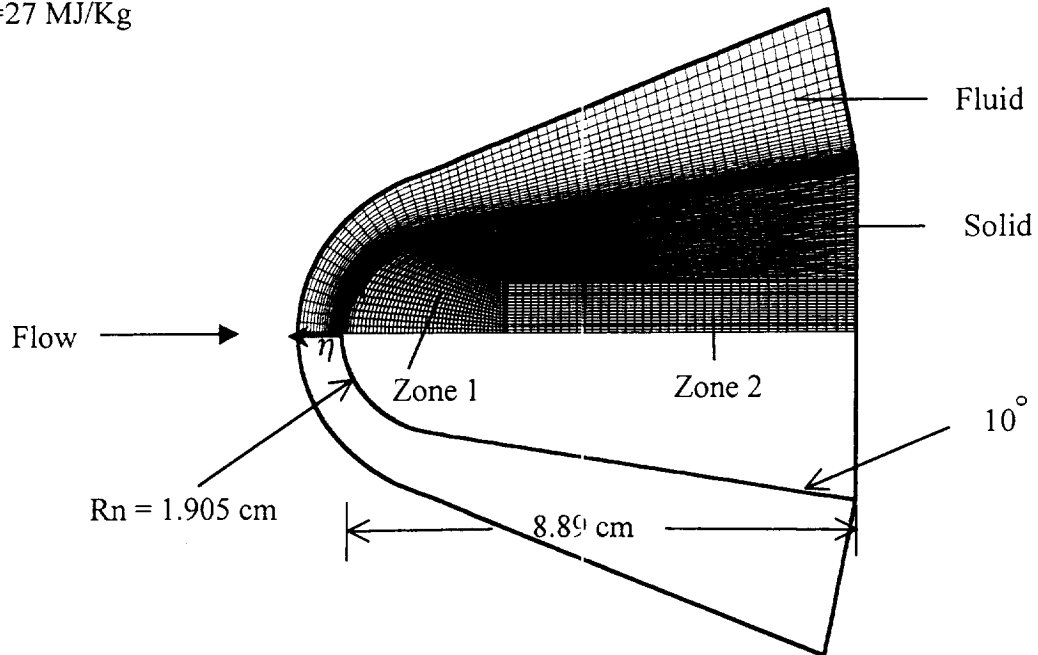


Figure 2: Computational grid for test series 1

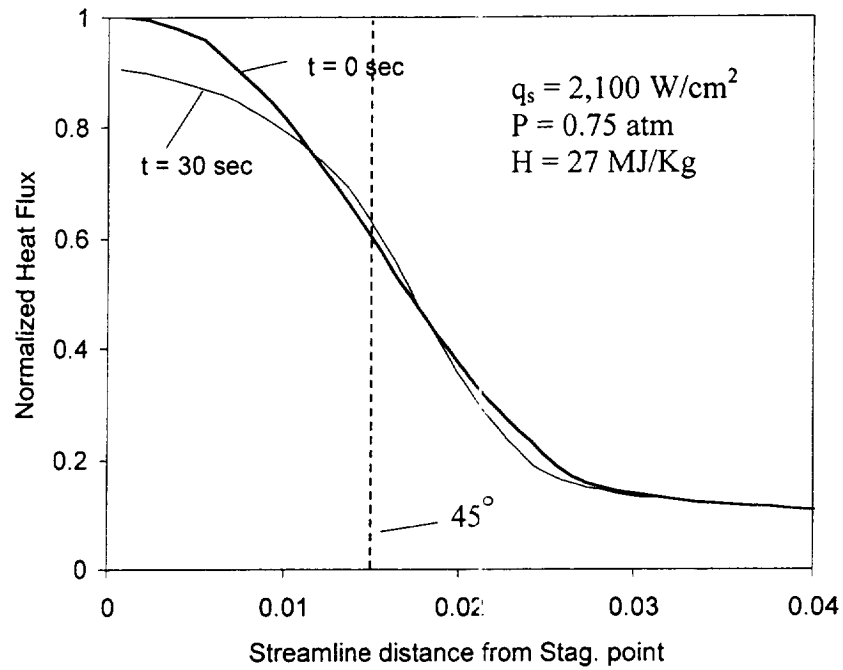


Figure 3: Heat flux distributions at $t = 0$ and 30 sec.

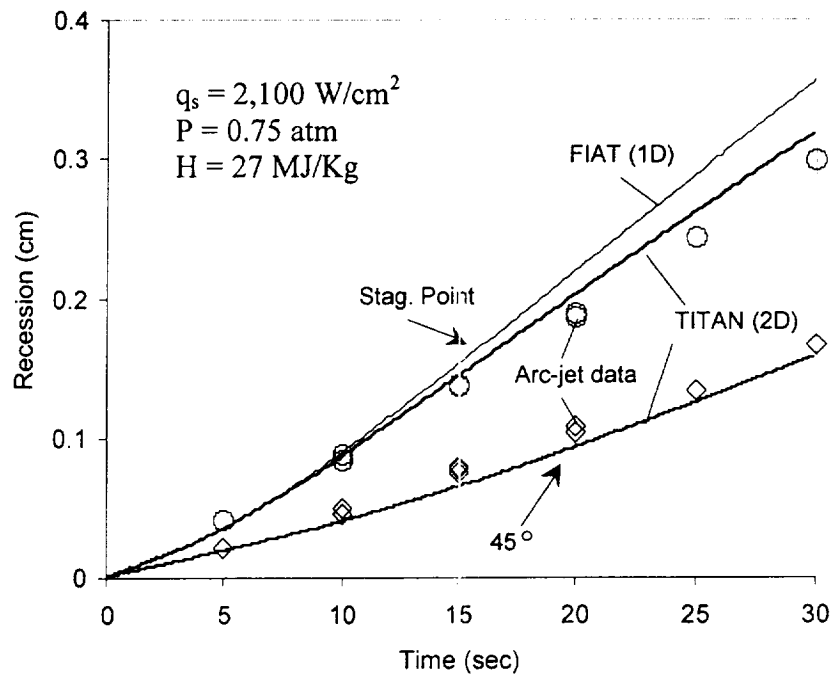


Figure 4: Computed and measured surface recessions

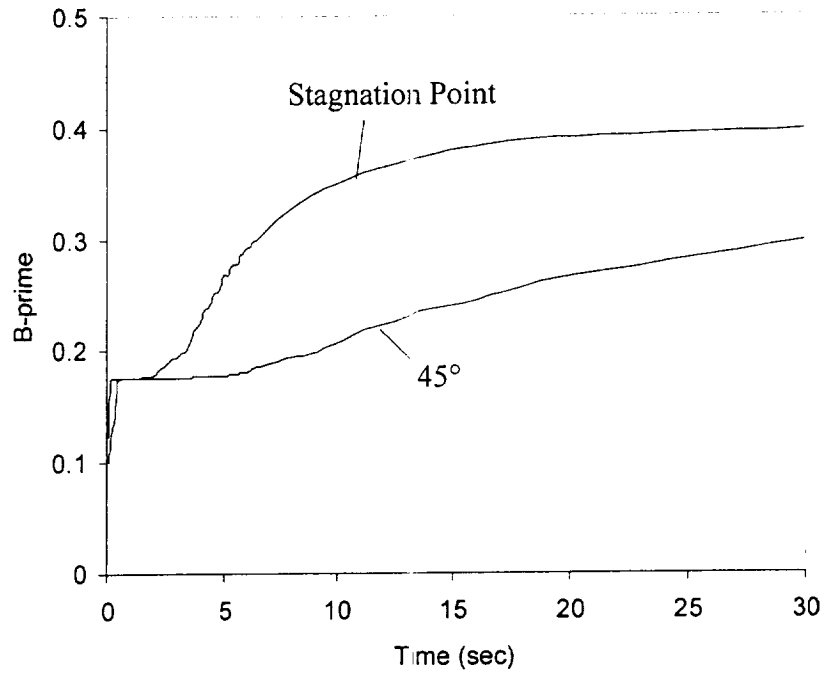


Figure 5: B' vs. time at two surface locations

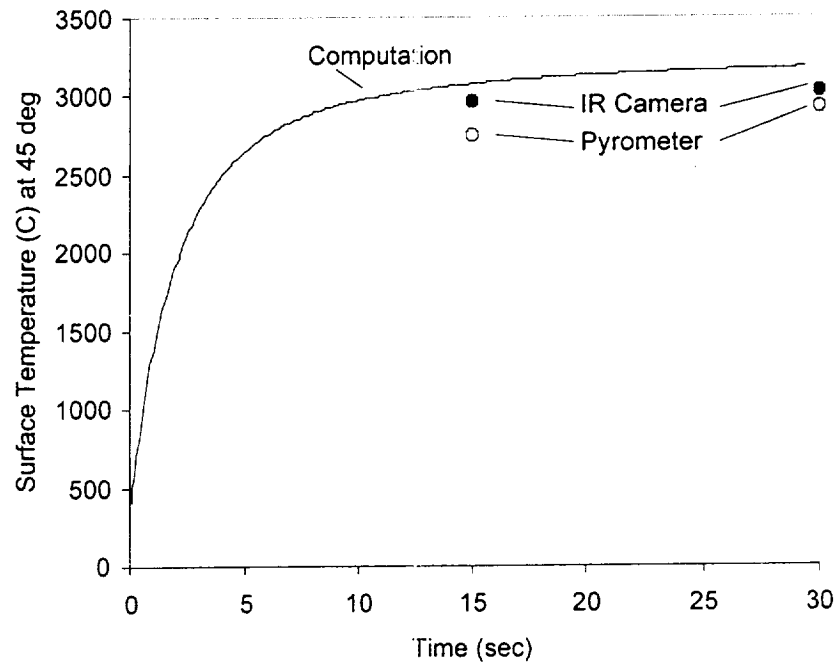


Figure 6: Surface temperature history at 45° from Stagnation Point

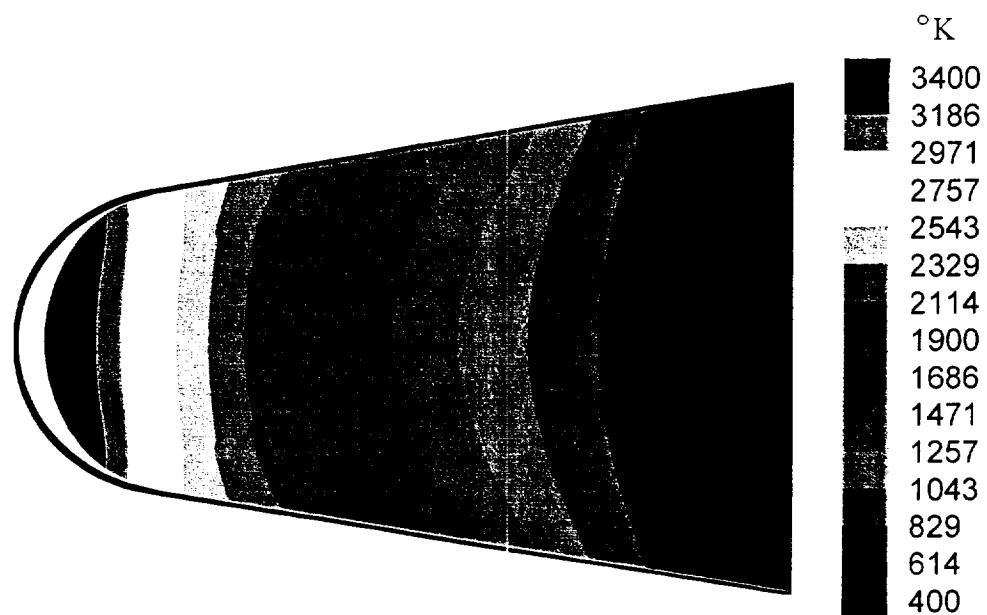


Figure 7: Temperature contours for graphite model at $t = 30$ sec

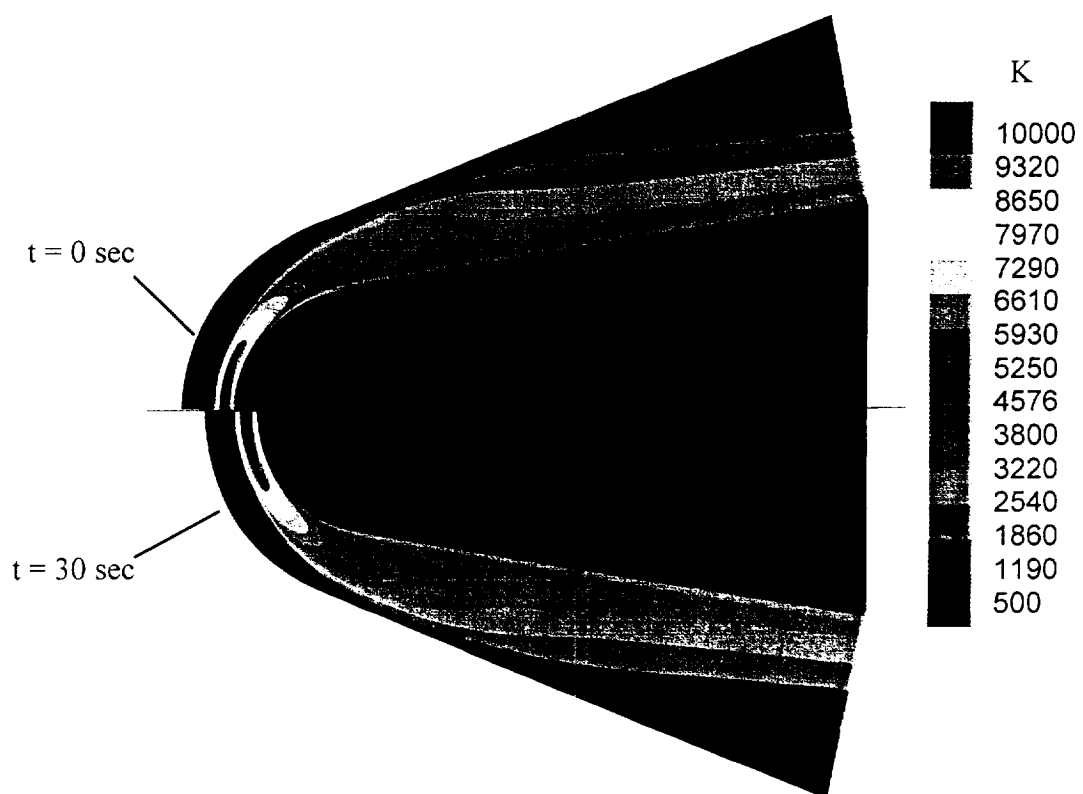


Figure 8: Flowfield temperature contours at $t = 0$ and 30 sec.

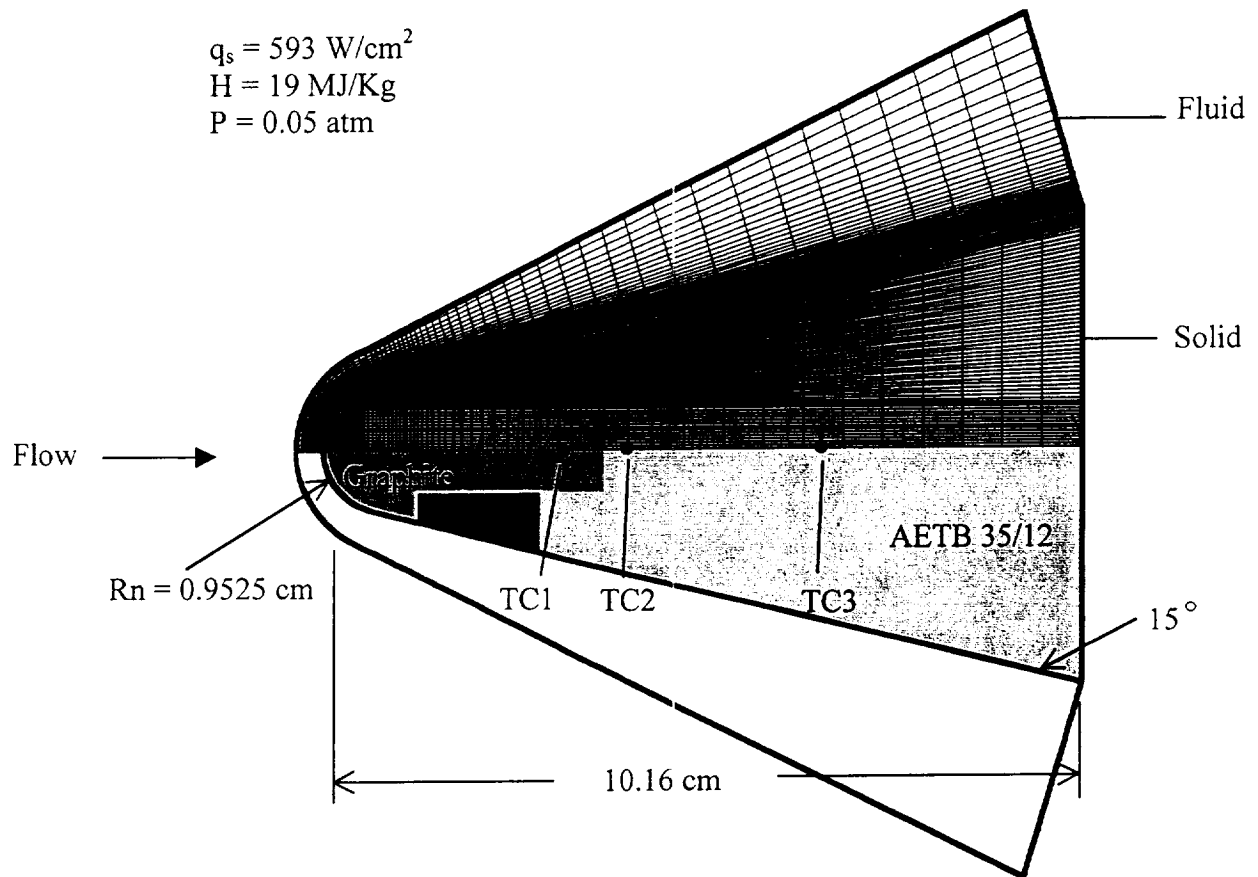


Figure 9: Computational grid for test series 2

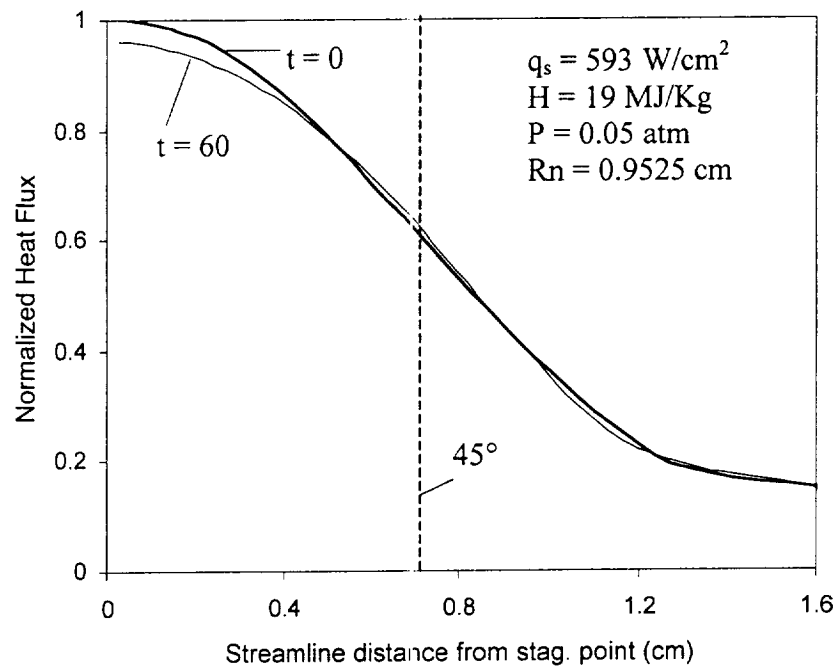


Figure 10: Heat flux distributions at $t = 0$ and 60 sec.

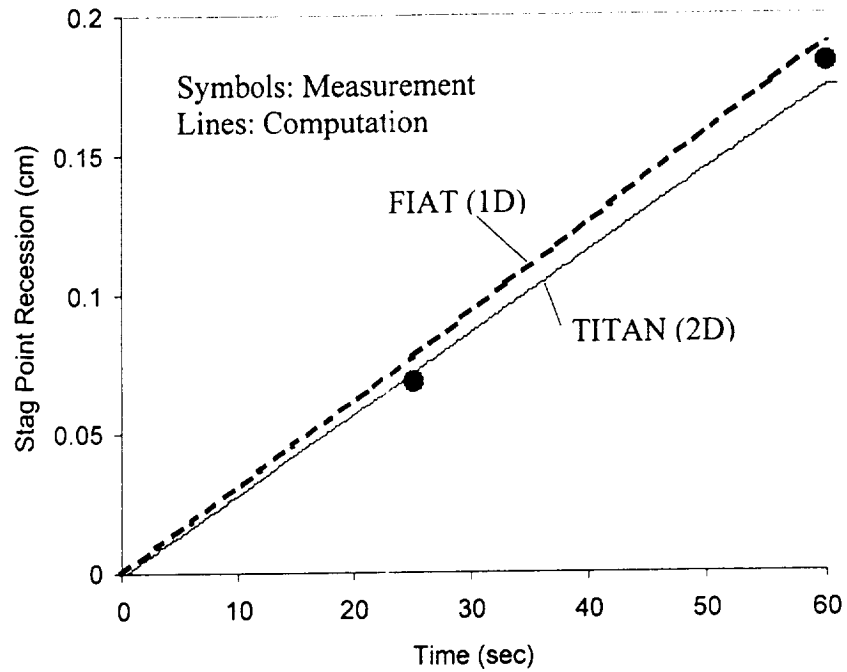


Figure 11: Computed and measured surface recessions

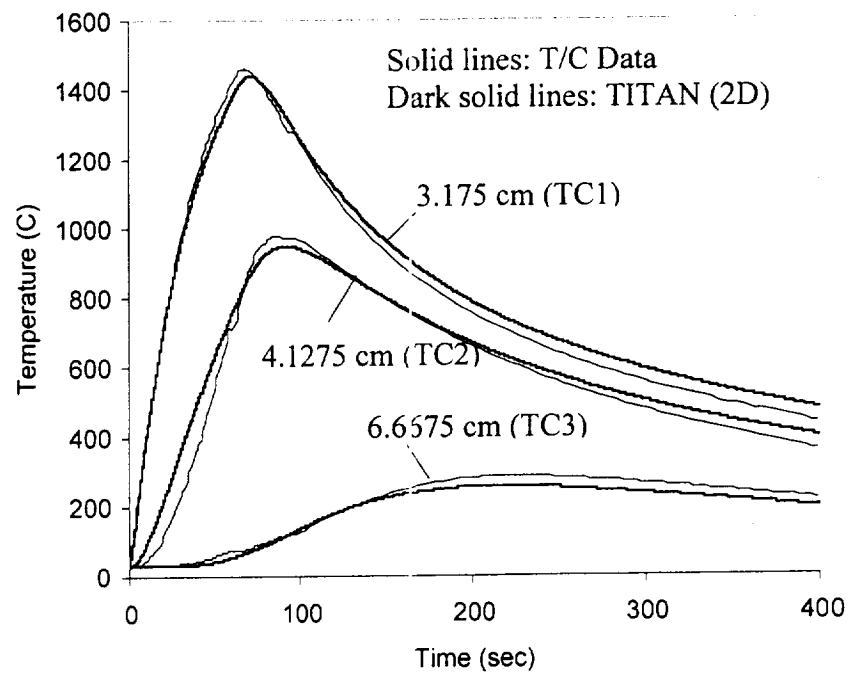


Figure 12: In-depth temperature histories along axis at three depths

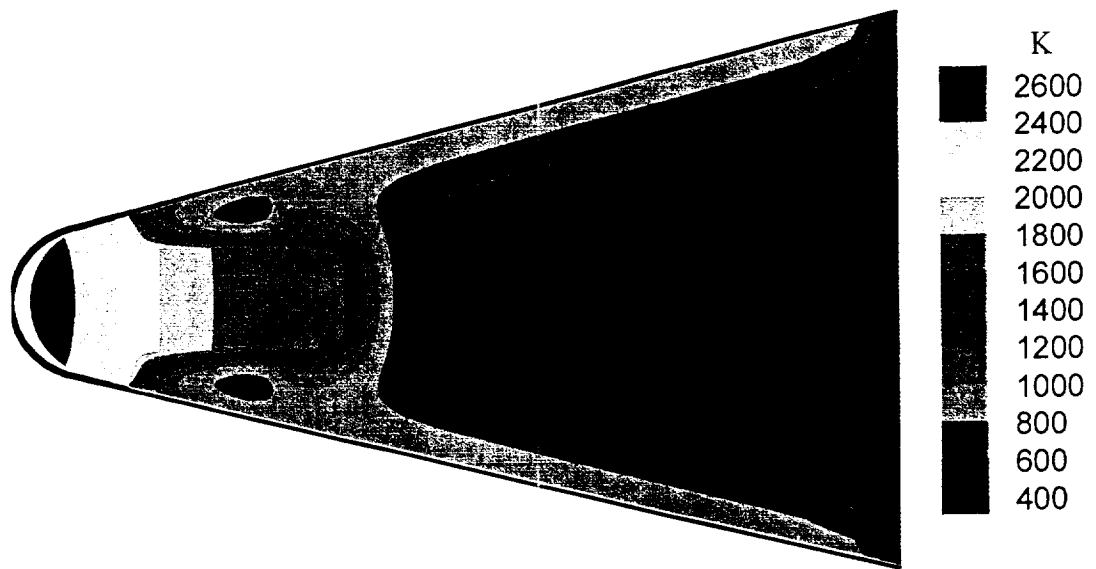


Figure 13: Temperature contours for composite model at $t = 60$ sec

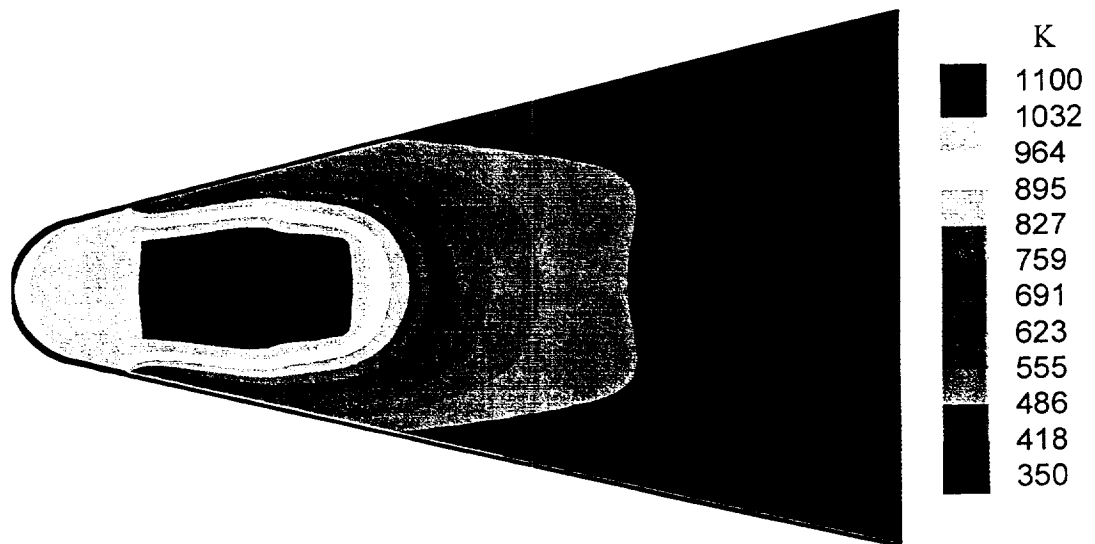


Figure 14: Temperature contours for composite model at $t = 200$ sec

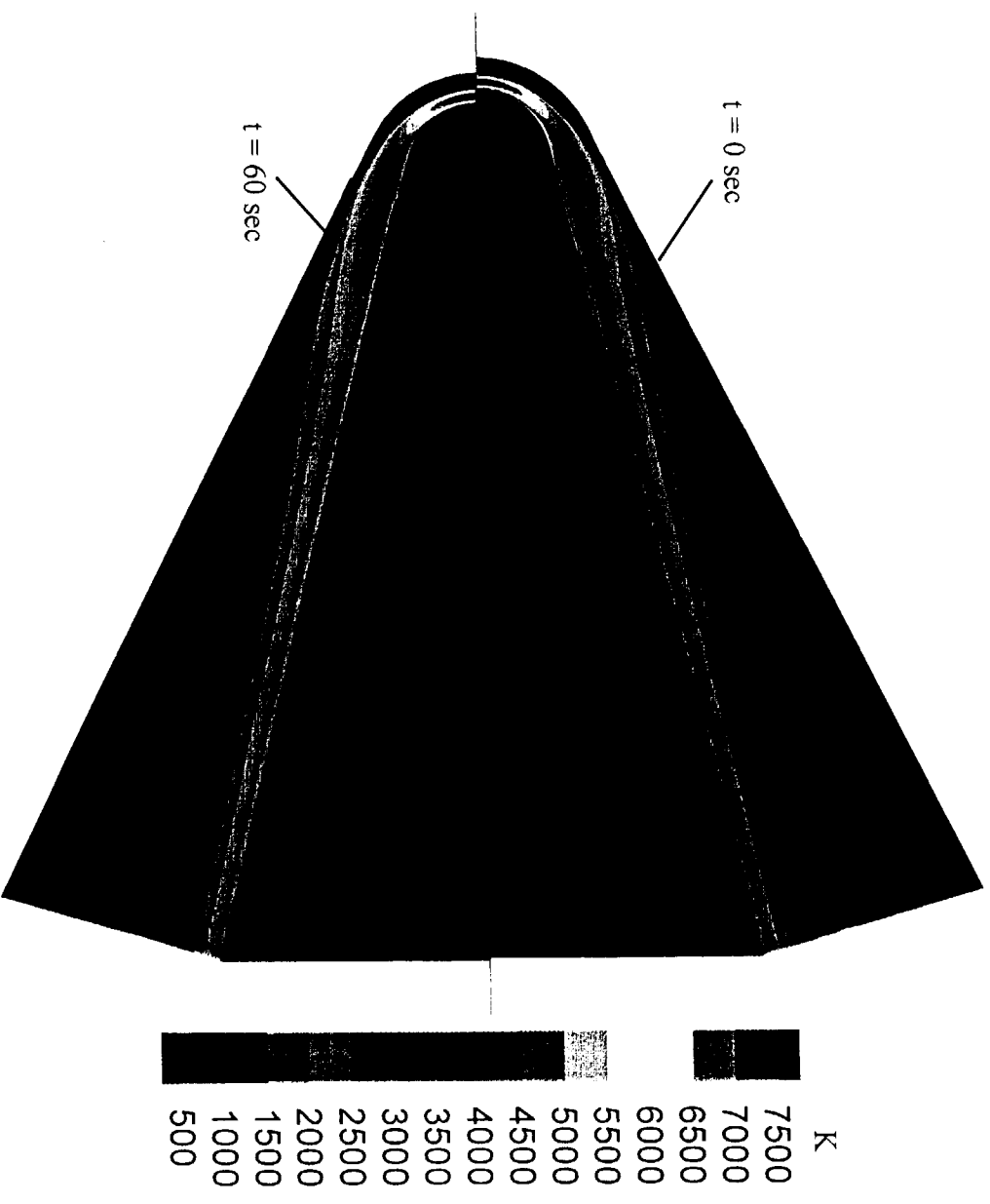


Figure 15: Flowfield temperature contours at $t = 0$ and 60 sec.